

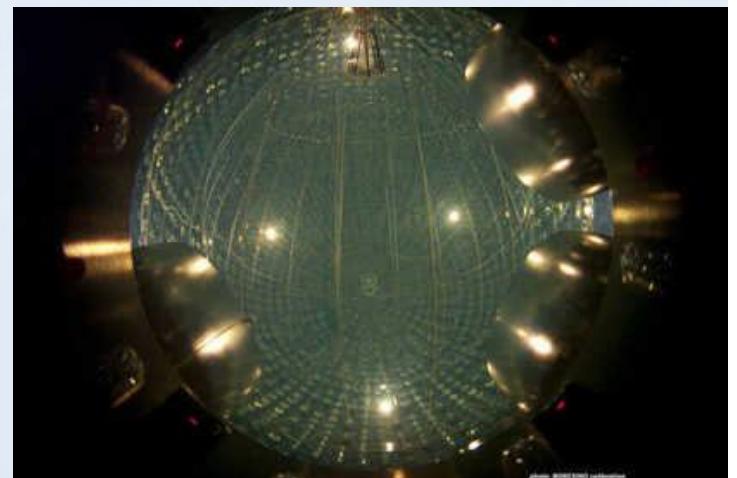
Reactor antineutrinos: update

B. Ricci for the ANTINU WG,
with help of F.Mantovani and S. Chubakov
BOREXINO general meeting 1-3 Dec. 2011, Milano

Outline



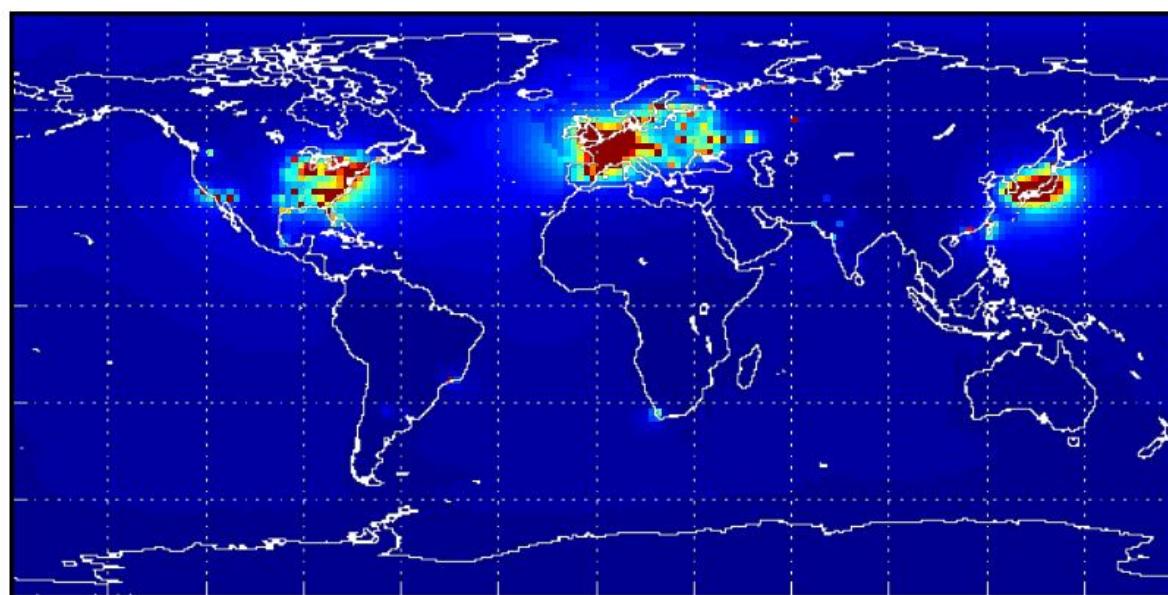
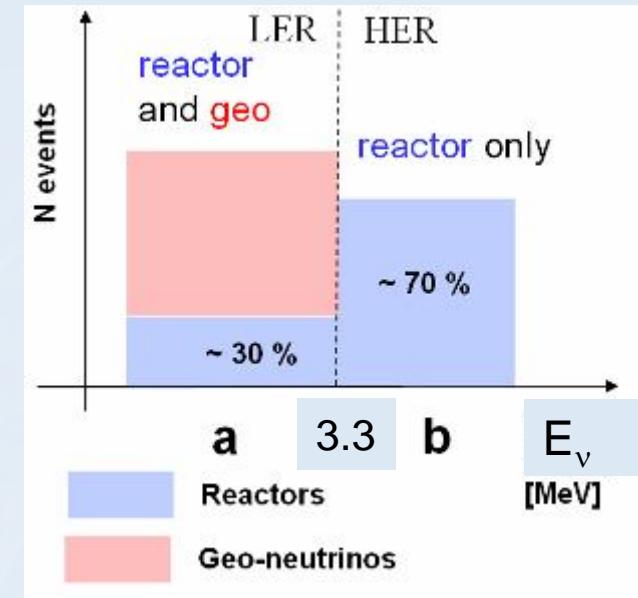
- why reactor anti-neutrinos ?
- reactors in the world
- signal calculation
- updates
- conclusion



Why reactor antineutrinos?

- Reactor antineutrinos are the main source of background in geo neutrinos detection

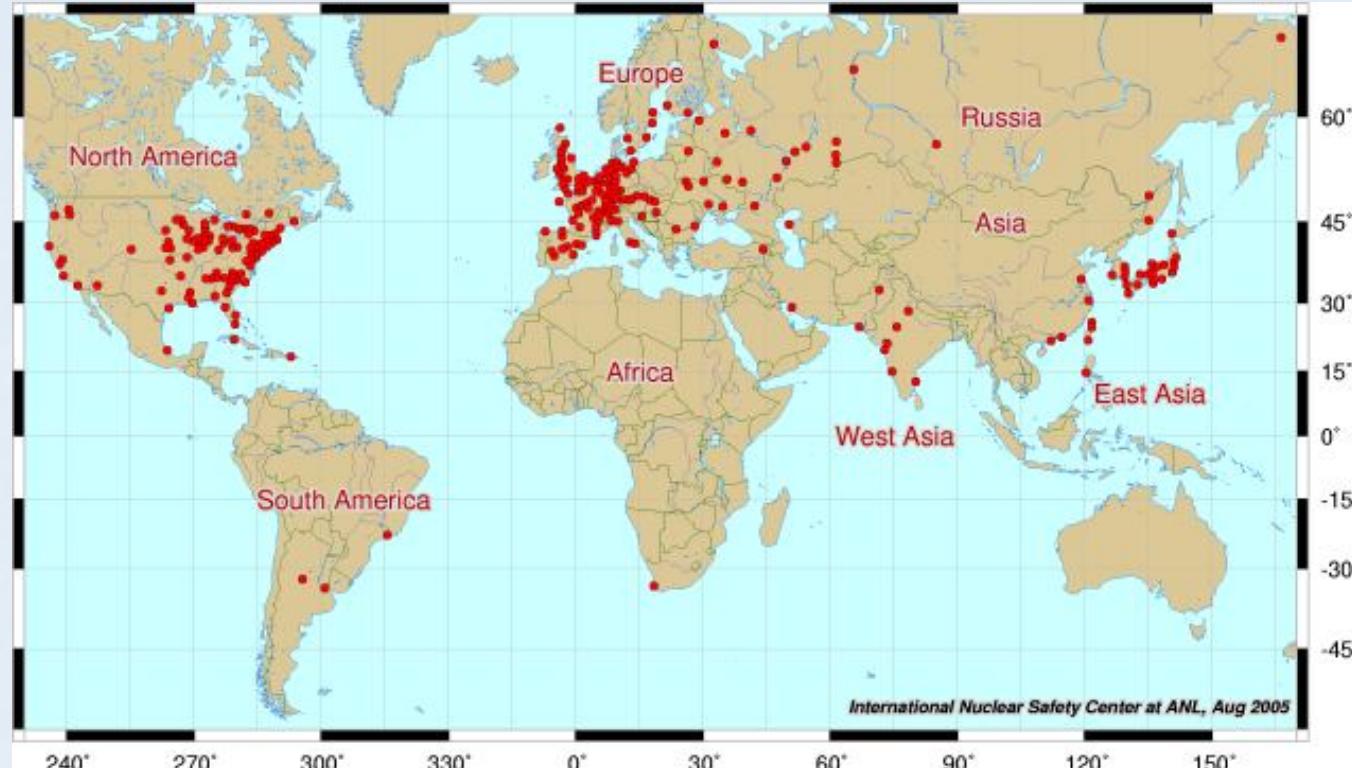
The map below* is based on 2000 IAEA database and considering all reactors at full power. The ratio r is referred to the geo-neutrino energy window.



	r
Kamioka	6.7
Sudbury	1.1
Gran Sasso	0.9
Pyhäsalmi	0.5
Baksan	0.2
Homestake	0.2
Hawaii	0.1
Curacao	0.1

*Fiorentini et al - Earth Moon Planets - 2006

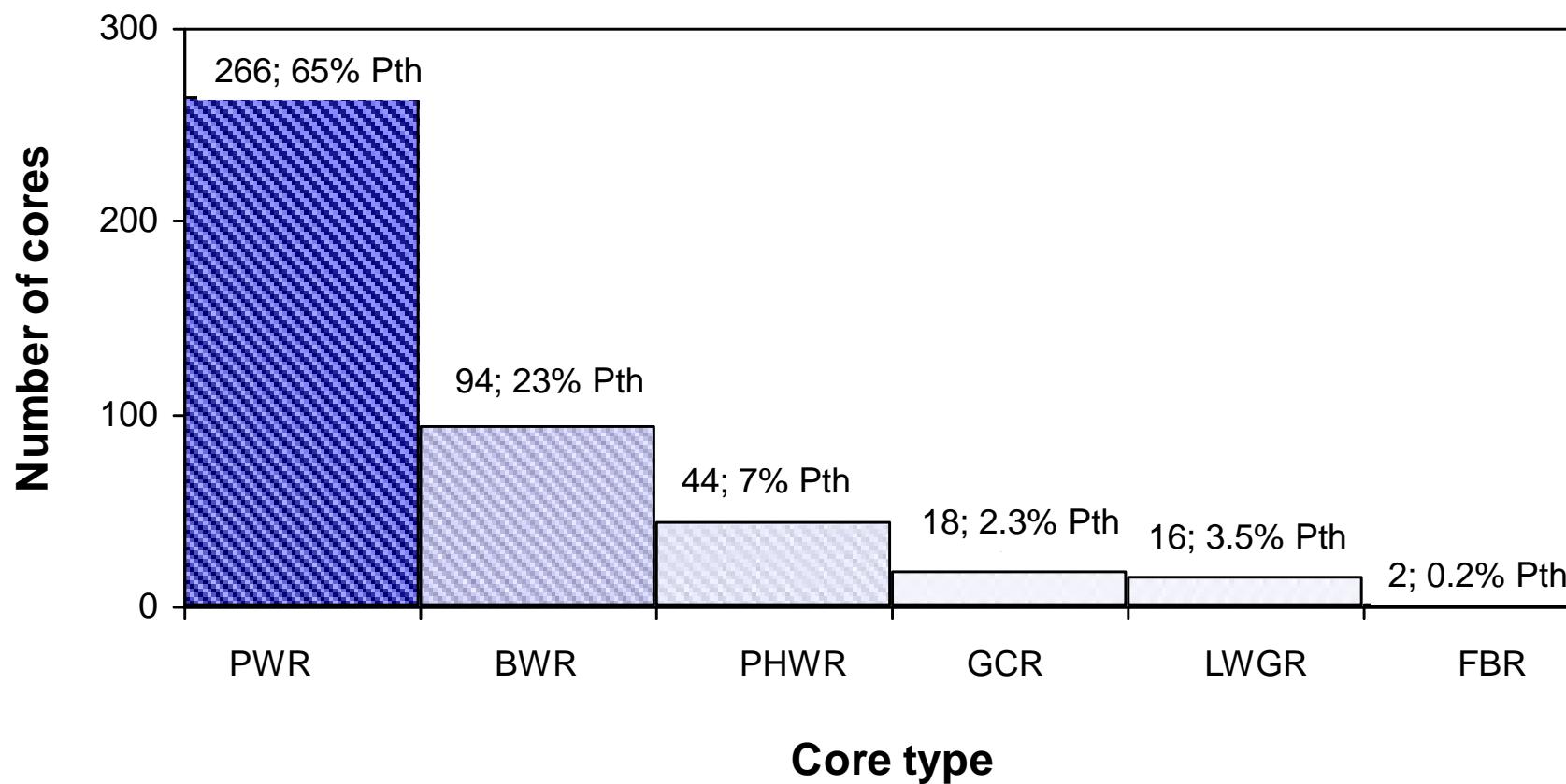
Nuclear power plants in the world



	#cores	Pth [GW]
■ Europe + Russia	197	519
■ North America	122	353
■ Japan+ Korea	76	201
■ Others	45	75
■ Total:	440	1148
■ Mean thermal power for core: 2.6 GWth		

at 31 Dec. 2009

Reactors by type



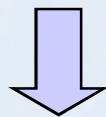
PWR	Pressurized (light) Water Reactor
BWR	Boiling Water Reactor
PHWR	Pressurized Heavy Water Reactor

GCR	Gas Cooled Reactor
LWGR	Light Water Graphite mod.
FBR	Fast Breeder Reactor

Signal calculation

DETECTOR

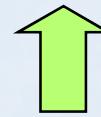
- $\varepsilon = 100\%$ detection efficiency
- $\tau = 1$ year
- $N_p = 10^{32}$



- ν PHYSICS
- P_{ee} = survival probability
 - $\sigma(E)$ = cross section
 - $\text{anti-}\nu_e + p \rightarrow e^+ + n$
 - $E_{th} = 1.806$ MeV
 - (calculation from Vissani and Strumia 2003)



$$N_{TOT} = \varepsilon \ N_p \ \tau \sum_{i=1}^{N_{reactor}} \frac{P_i}{4\pi d_i^2} <LF_i> \int dE_\nu \sum_{k=1}^{N_{fuel}} \frac{p_k}{Q_k} \lambda_k(E_\nu) P_{ee}(E_\nu, d_i) \ \sigma(E_\nu)$$



REACTOR

- d_i = reactor distance
- P_i = reference thermal power
- LF = Load Factor
- p_k = power fraction

NUC. PHYS.

- Q_k = energy released for fission
- λ_k = reactor anti-neutrino spectrum

K=235U, 238U, 239Pu, 241Pu

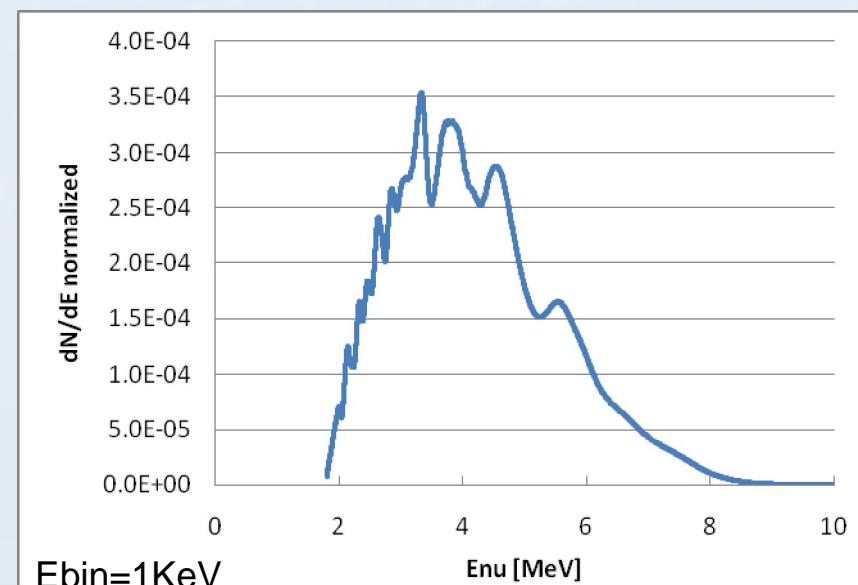
Starting point: March 2010

- vacuum oscill, 2 neutrinos
- “standard” PWR fuel + 35 mox
- LF : 1 Jan 2008 - 31 Dec 2008
- spectrum:Huber &Schwetz 2004
- 100% detection efficiency
- $E_{\nu} = 1.8\text{-}10 \text{ MeV}$

$$S = 93.3 (1 \pm 5.4\%) \text{ TNU}$$

$\Delta m^2 = 7.65 \times 10^{-5} \text{ eV}^2$
$\theta_{12} = 33.46^\circ$

Power Fraction				
	235U	238U	239Pu	241Pu
PWR	0.56	0.08	0.3	0.06
MOX	0	0.08	0.708	0.212



Update 1: LF for 2009-2010

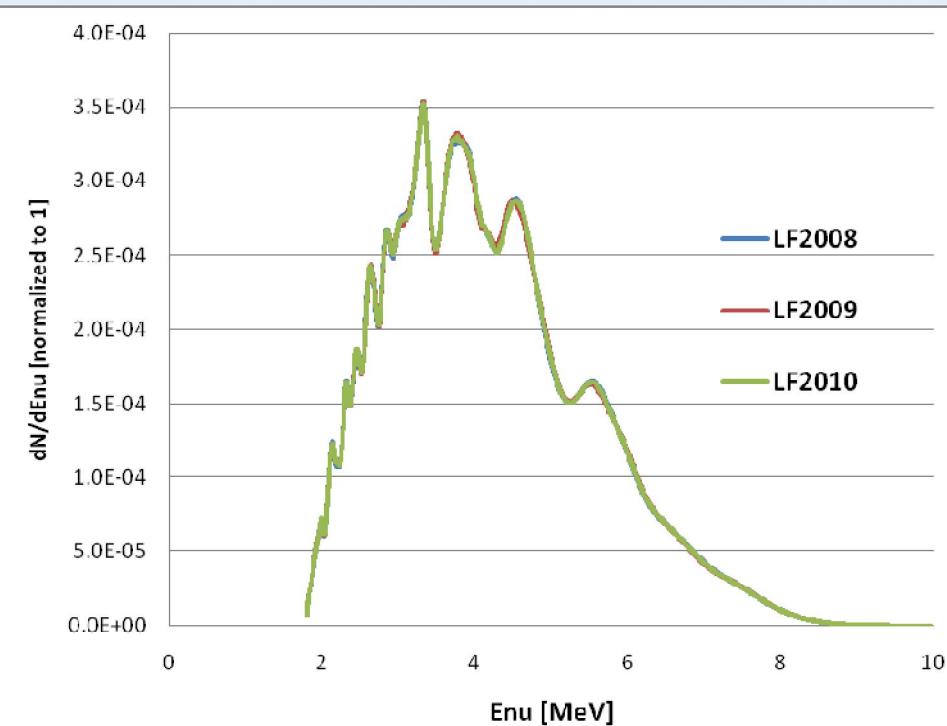
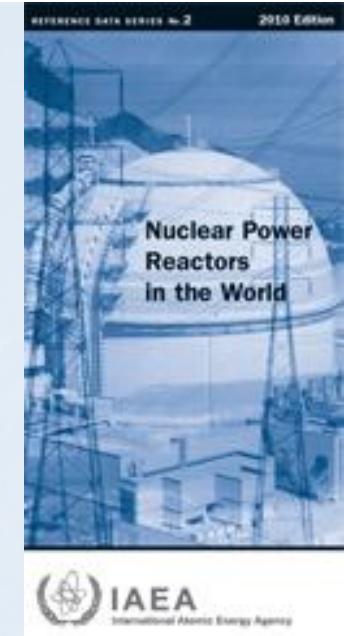
- Expected Signal at LNGS (100% effic.) by using LF of different years (weighted average with days in each months):

LF2008 93.3 TNU

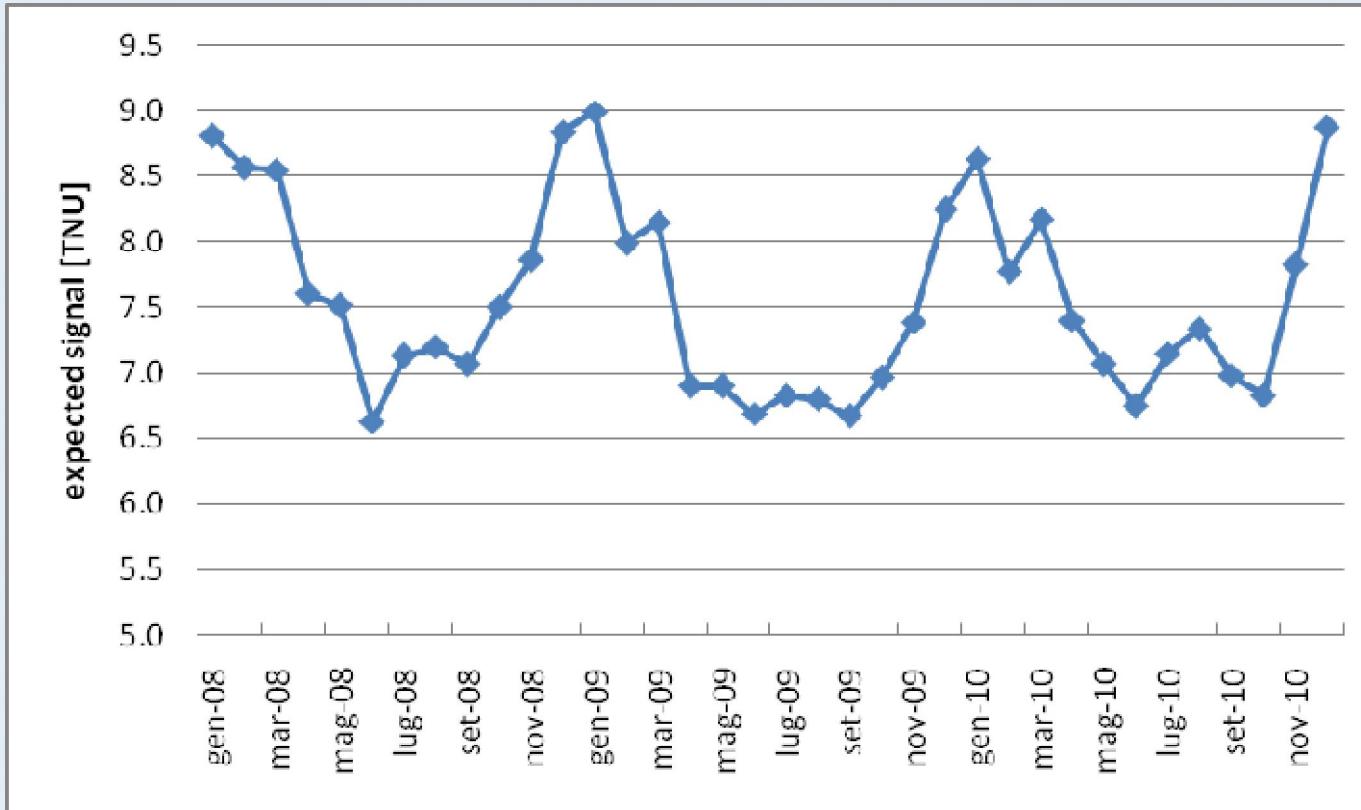
LF2009 88.6 TNU

LF2010 90.8 TNU

av. value: 90.9 TNU



LF and time variation



- about 30% variation (max-min)
- few events for month

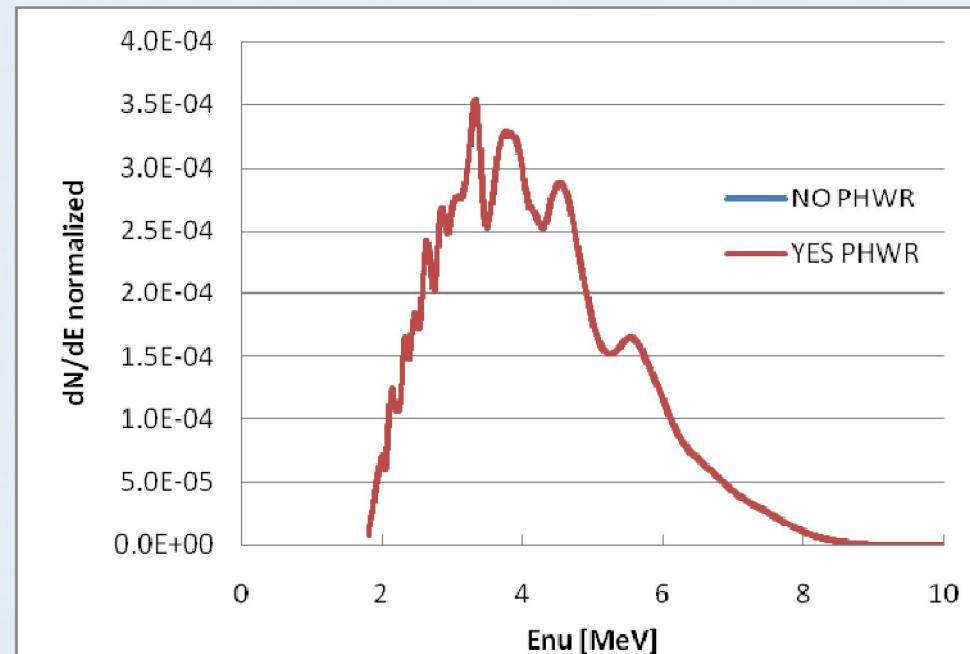
-100% efficiency
- $N_p = 10^{32}$ protons

Update 2: PHWR fuel

	235U	238U	239Pu	241Pu
PWR	0.56	0.08	0.3	0.06
PHWR	0.54	0.41	0.02	0.024

- In 2010, 46 PHWR cores in the world
(2 cores in Romania)

- No effect on signal:
NO PHWR 90.9 TNU
YES PHWR 90.9 TNU

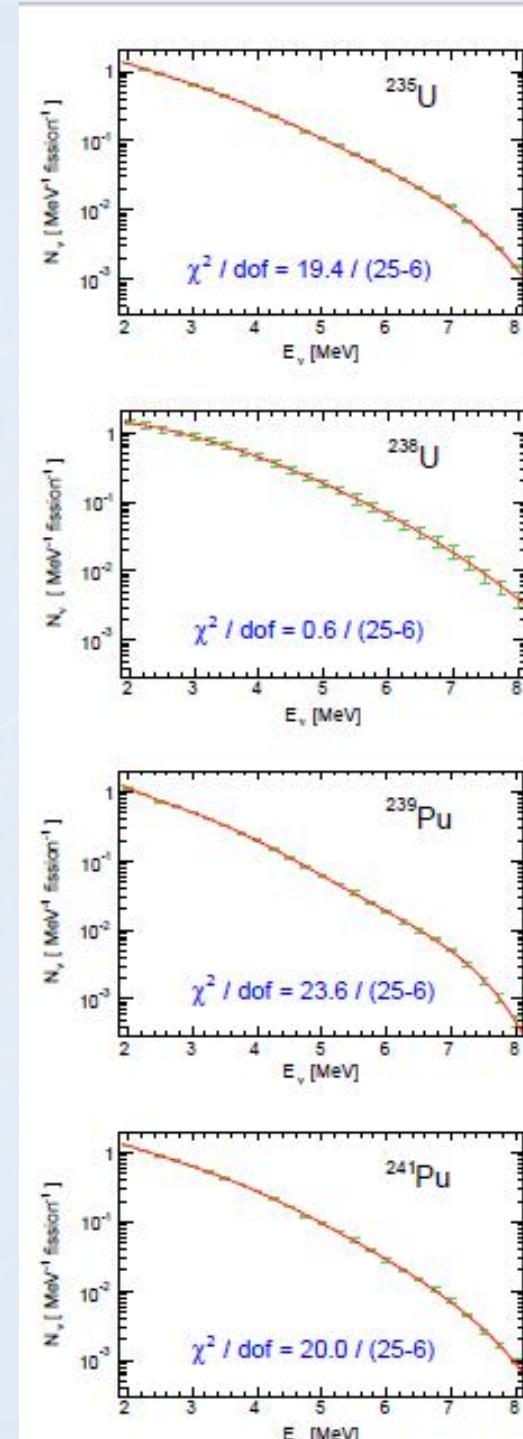


Update 3: reactor antineutrino spectrum

1) Huber and Schwetz 2004
(used in geo nu -paper)

2) Mueller et al. 2011

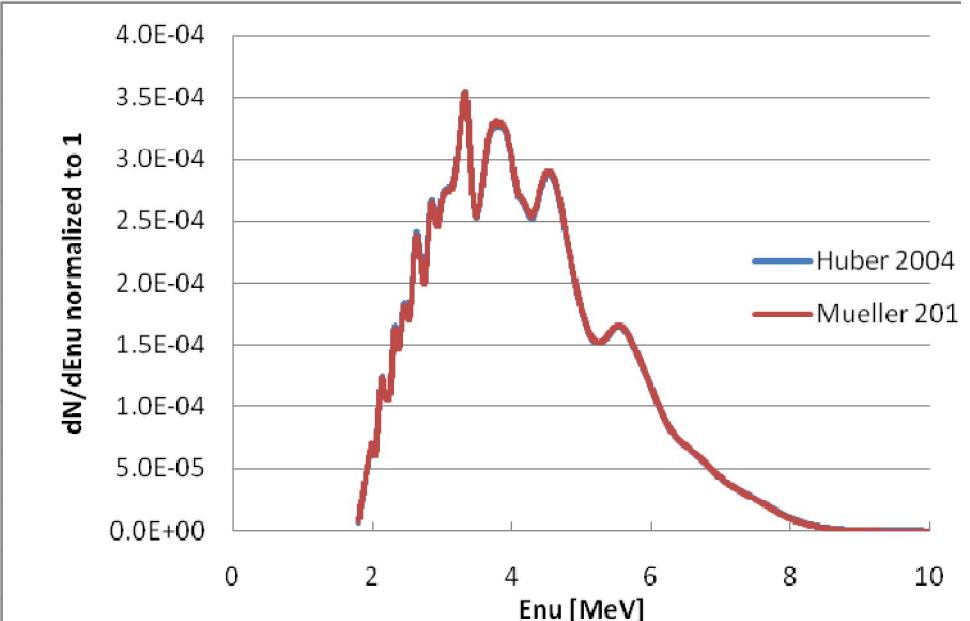
“...While the shapes of the spectra and their uncertainties are comparable to that of the previous analysis ..., the normalization is shifted by about +3% on average...”



Antinu spectrum and signal

- The predicted signal at LNGS, 100% efficiency, averaged on 3 years is:

Huber 2004: 90.9 TNU
Mueller 2011: 94.1 TNU



- By using Mueller 2011 the signal increases (+ 3.5%)

Conclusion

- By using:
 - $\langle LF \rangle$ in 3 years (2008-2009-2010)
 - PHWR fuel
 - reactor antinu spectra from Mueller et al 2011
 - 100% detection efficiency
 - $E_{\nu} = [1.8 - 10 \text{ MeV}]$
 - Oscillation in matter : +0.6%
 - Spent fuel: + 1%

One has: $S = 95.6 (1 \pm 5.4\%) \text{ TNU}$

i.e.:

$$N_{\text{reactor}} = 17.2 (1 \pm 5.4\%) \text{ events/300ton/yr}$$

$$\begin{aligned} N_{\text{protons}} &= 6.02 \times 10^{28} / \text{ton} \\ \text{yr} &= 365 \text{ d} \end{aligned}$$