

AMBER Status Review

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Meeting with spokespersons on June 8, 2021

142nd SPSC Meeting, June 11 2021

AMBER Collaboration

News and Progress in Formation of the AMBER Collaboration

- Election and nomination of
 - Physics Program Coordinators
 - CB chair
 - Technical Board
 - Spokesperson election just finished. Oleg Denisov and Jan Friedrich in charge for the initial mandate as AMBER spokespersons
- Collaboration name fixed to **AMBER** (resp. NA66/AMBER)
- MoU drafting ongoing since 15 April (bi-weekly) Chaired by Andrea Bressan
- First Collaboration Meeting took place May 6 and 7
- Stimulating two-days meeting with about 80 participants
- Current collaboration strength 189 (voting members)

Caveat: not all the funding to start the physics run has been secured yet

AMBER Mission





In order to "prove" that QCD underlies the EHM phenomenon to compare Lattice and Continuum QCD calculations with experimental data by measuring:

- 1. Quark and Gluon PDFs of the pion/kaon/proton
- 2. Hadron's radii (confinement)
- 3. Excited-meson spectra

AMBER Long Range Plan



Proton Radius Measurement Antimatter production cross section Pion structure (PDFs) via DY and charmonia

Phase-1 Proposal approved by RB on 02/12/2020

Kaon and pion structure (PDFs and PDAs) High precision strange-meson spectrum Kaon and pion charge radius Kaon induced Primakoff reaction

Phase-2 Proposal submission in the beginning of 2022

Proton Radius



Uncertainty on G1





statistical precision of the proposed measurement, down to $Q^2 = 0,001 \text{ GeV}^2/c^2$, Cross section is normalised to the G_D - dipole form factor

AMBER:

Pure muon beam + magnetic form factor contribution suppression at high E

The impact of the proposed p + p measurements on constraining the production of cosmic anti-protons versus their kinetic energy. Each curve represents the fraction of anti-proton production phase space as constrained by AMBER cross section measurements in p-p, p-He and He-p channels, compared to NA61 (p-p) and LHCb (p-He) measurements



10/06/2021

Pion Structure

11/June/2021

AMBER Pilot Run

Main Goal to be reached: Proof-of-principle of all new detector equipment

- Test the IKAR TPC in dedicated 20 days of beam (CEDAR position)
 - determine the noise/background induced by the muon beam, detect proton recoils correlated with scattered muons
- Test of the unified tracking detector station
 - operate one detector station with prototypes of both the silicon-pixel detector and the scintillatingfibre hodoscope
- Test of the new DAQ system (possible for TPC in park position)
 - operate new free-running DAQ system for readout for all new detector components

AMBER Pilot Run

Need to be complemented with a tracking system for the scattered muon

IKAR TPC (currently at CERN):

- \rightarrow 2x drift cells with 400 mm length
- \rightarrow New adapted field-shaping rings
- \rightarrow Anode structure: identical structure, but with smaller diameter wrt final TPC
- \rightarrow Operation pressure of max. 10 bar
- \rightarrow New power-supplies and front-end electronics
- Ready for first pressure test
- Dedicated gas-purification system will be used

Licence issue with ALPIDE detectors Negotiation pursued by ALICE Likely not ready for the pilot run Parasitic test assumed

Combined Silicon-Fibre tracking station:

- ightarrow 2x+2y planes of 500 μ m scintillating fibre 9.6x9.6 cm²
- ightarrow 3 planes of pixel-silicons 9x9 cm² (pixel size 28x28 μ m)
- \rightarrow Operation pressure of max. 10 bar
- \rightarrow New power-supplies and front-end electronics
- Small distance between the Silicon-pixel detectors and the Scintillating-Fibre Hodoscope
- · Allow for independent access and cooling infrastructure
- Compatible with beam line elements for the He volume

Pilot Run Tracking

Several options available for tracking available

Tracking required to evaluate TPC response in the highintensity muon beam and correlation of events.

- Option 1: ALPIDE telescope:
 - $\rightarrow\,$ Same technology as in the main run
 - \rightarrow 9-plane telescopes are available
 - $\rightarrow\,$ Dedicated DAQ system based on EUDAQ
 - → Requires hit-time information due to timeresolution or delicate tracking algorithm
- <u>Option 2:</u> COMPASS Silicon-strip detectors:
 - \rightarrow Comparable hardware as in 2018
 - → Based on existing COMPASS DAQ, event definition and reconstruction
 - $\rightarrow\,$ No additional timing detectors required
- Under investigation: Possibility of an additional "physics" trigger to enrich correlated events

Both options are under discussion and evaluation to find the best option in terms of reconstruction, time schedule and man power

Pilot Run Tracking

Usage of silicon-strip tracking detectors

Previous model of the new cryogenic COMPASS silicon-strip tracking detector from 2004.

- Used in test measurement in 2018
- Characteristics:
 - \rightarrow 300 µm wafer with an active area of 70x50 mm²
 - → 1280 (1024) readout strips per projection and utilisation of charge sharing between strips
 - \rightarrow Four single projections: X,Y and U,V (rotated by 5°)
 - → Time resolution: σ = 1.5 ns, spatial resolution: σ = 5 10 um (depending on cluster size)
 - → Triggered readout based on APV25 chips
- Disadvantages:
 - No experience with ALPIDE / tracking for main run
 - Trigger rate limited at 40 kHz
- Advantages:
 - Fully available and working system
- Todo:
 - "Minor" work: check equipment and prepare spares etc.

No cryogenic but part of the services from in-operation COMPASS silicon

11/June/2021

142nd SPSC Meeting

Trigger and DAQ

TRLO

Synchronisation using time stamps

Events from separate TPC and silicon DAQ tagged with time stamp.

- TRigger LOgic (TRLO) time stamp with 100 MHz resulting in a time resolution of 10 ns
- Synchronisation of time stamps during SPS spill

- Linear interpolation between synchronisation time stamps for calibration
- Match TPC and silicon events via TRLO time

Trigger_hitPositionY:Trigger_hitPositionX {Trigger_detectorId==2302}

11/June/2021

Beam Optics Study

Beam Optics - Focused Beam (M2 Test Beam Area)

Beam Optics - Parallel Beam (M2 Test Beam Area)

Minutes

139th SPSC Meeting:

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The Committee **recommends** approval of the proposal SPSC-P-360 by the AMBER Proto-Collaboration to use the M2 beam-line before LS3 to perform measurements related to:

- (i) Drell-Yan and J/Psi production using the conventional M2 hadron beam;
- (ii) proton-induced antiproton production cross sections for dark matter searches;
- (iii) the proton charge radius using muon-proton elastic scattering.

The proton-radius program is contingent on a successful pilot run previously approved for the first year of SPS operation after the Long Shutdown LS2.

The SPSC **emphasises** that the beam time allocated to AMBER for any of the measurements is subject to the available time, and to annual discussions in the Committee on the use and sharing of the M2 beam line.

Suggested minutes of the 142th SPSC meeting:

The SPSC **recognizes** the physics interest of the AMBER program and **is looking forward** to the consolidation of the Collaboration structure.

The SPSC **aknowledges** the progress in preparation of the AMBER pilot run dedicated to the validation of the proton radius measurement and encourages the Collaboration to concentrate on the most viable option in complementing the TPC tracking.

M2 Schedule in 2021

SPS user schedule for 2021															CERN										
schedule issue date: 04-Jun-2021					Version: 2.1					Exp. PS/SPS Exp. O					ther Exp. INT Exp.				0.1						
	Week		Jun		27	Jul 28	Jul		31	32	33	33 34		35 36 3		, 38 3		40	Oct		43	A3 44 45			
Machine				20	29	50	51	52					TS1 Cold	ex		+0	41	42			45	RF			
T4 - K12		SF T So			NA Setup 7									N / 1	462 26										
T6 - M2				SPS & TT20 Setup 7	NA Setup 7				NA58 COMPAS 86				ass	5					Amber 21			NA64 mu 19			

COMPASS minimum run to justify a meaninful data-taking

(3 weeks for spectrometer commissioning assumed to happen in July anyway)

2 weeks for target commissioning

2 weeks for each target polarization cycle (x3 minimum)

ends up to a bare minimum of 8 weeks = 56 days

AMBER and NA64mu pilot runs are crucial for the timely start of their RUN3 physics program

A beam schedule change request is possible but too premature to be considered at this time COMPASS and AMBER Collaborations largely overlap and might eventually negotiate a compromise A request for beam time compensation might impact 2022 beam schedule.

Not clear SPSC role if a decision has to be taken before fall meeting

AMBER Vision

O. Denissov presentation at the open session:

Starting point depends of the semiconductors availability on the market

SPSC Discussion in October '20

2022 schedule still in jeopardy

Pending information on readiness from COMPASS (polarized taregt), AMBER (ALPIDE telescope) MuonE (tracking stations), on potential 3 weeks start delay and Pb ion run length....

Need to keep trace with the experimetn for an extensive discuiion in fall meeting