## Upcoming at Jlab up to 2027-2028 (Hall A and Hall A -> Hall C)

G.M. Urciuoli

- Neutron Star structure (Hypernuclear spectroscopy)
- Moller experiment

## Why Study Neutron Stars?

Frontiers in particle, nuclear, condensed matter, plasma, hydro physics, QED, and general relativity

Core composition is unknown: Nucleons? Hyperons? Quark matter?



Matter in NS cores reaches several times nuclear density

It is very neutron rich, unlike nuclei

It is also very cold:  $kT << E_F$ T~10<sup>9</sup> K is cold!

From Cole Miller's Talk on PREX/CREX and Astronomical Observations of Neutron Stars at 2023 Summer Hall A/C Meeting

http://www.astroscu.unam.mx/neutrones/NS-picture/NStar/NStar-I.gif

## NS masses

- A given equation of state (EOS) P(ε) (P is pressure, ε is total mass-energy density) predicts M(R) Assume equilibrium
- Also predicts maximum mass
- Viable EOS must accommodate largest measured mass



## Double NS Masses

- Very tightly clustered M=1.35+-0.1 M<sub>sun</sub>
- Does this indicate a very low upper limit on masses?
- Or are formation conditions just similar?



http://www.lsw.uni-heidelberg.de/users/mcamenzi/NS\_Mass.jpg

- $\sim 2 M_{sun}$  Neutron Stars
  - J1614–2230, 1.908+–0.016
     Demorest et al. 2010
  - J0348+0432, 2.01+-0.04 M<sub>sun</sub> Antoniadis et al. 2013
  - J0740+6620, 2.08+-0.07
     Cromartie et al. 2019
  - Eliminate EOS that are too soft, i.e., whose pressure is too low at the relevant densities



The Shapiro delay is a general-relativistic increase in light travel time through the curved space-time near a massive body. For highly inclined (nearly edge-on) binary millisecond radio pulsar systems, this effect allows us to infer the masses of both the neutron star and its binary companion to high precision. Here we present radio timing observations of the binary millisecond pulsar J1614-2230 that show a strong Shapiro delay signature. We calculate the pulsar mass to be (1.97+/-0.04)M<sub>solar</sub>

Demorest et al. 2010

Hyperons are expected to appear in their core at  $r \sim (2-3)r_0$  when  $m_N$  is large enough to make conversion of N to Y energetically favorable. However, this results in a reduction of the Fermi pressure exerted by the baryons and a softening of the equation of state (EOS). As a consequence, the maximum mass determined by the equilibrium condition between gravitational and nuclear forces is reduced. Most of EOS of matter containing strangeness predict a maximum neutron star mass of about 1.5 solar masses. However, the recent measurements of neutron star masses as big as 2 solar masses require a much stiffer EOS (Hyperon puzzle).



The hypothesis that Neutron Star masses cannot be greater than 1.5 solar masses derived from the assumption that the attractive interaction between  $\Lambda$  and nucleons, experimentally confirmed only at the normal nuclear density ( $\rho_0$ ), holds true also when one deals with higher density nuclear matter.

To solve the hyperon puzzle it has been suggested that three body forces could provide additional repulsion making the EOS stiffer enough.

This hypothesis is justified by the observation that realistic *NN* interaction models, which describe all the *NN* scattering data and light nuclei (by adding attractive *NNN* 3BF), fail to reproduce the nuclear saturation density  $\rho_0$ . This indicates that a repulsive, short-range 3BF is present in the interaction between nucleons.

It is natural to hence to postulate a similar repulsive  $\Lambda NN$  (and  $\Lambda\Lambda N$ ,  $\Lambda\Lambda\Lambda$ ) interaction. However, this has to be proved experimentally!





D.Lonardoni et al., Phys. Rev. Lett. 114, 092301 (2015) (AFDMC)

### How can we solve the hyperon puzzle?

Why is the hypernuclear spectroscopy program at Jlab so important?

Why is very important to determine  ${}^{208}_{\Lambda}Tl$  binding energies in addition to  ${}^{40}_{\Lambda}K {}^{48}_{\Lambda}K$  ones?



From Tamura's HYP2022 Proceedings

### Precise $B_{\Lambda}$ values:

@Jlab:

- Precise binding measurement
- High resolution apparatus



Precise  $B_{\Lambda}$  values in a wide range of  $\Lambda$  hypernuclei



The spacing of the single-particle energies as a function of A put more constraints on the theoretical fits.



A satisfactory description of the hyperon separation energies in a wide mass range and for the  $\Lambda$  occupying different single particle state orbitals (s, p and d wave) has be obtained recently by auxiliary field diffusion Monte Carlo (AFDMC), using a phenomenological interaction in which the two-body potential has been fitted on the existing  $\Lambda$ p scattering data and with the inclusion of the three-body  $\Lambda$ NN force. However, these potential models predicting relatively small differences in the  $\Lambda$  separation energies of hypernuclei give dramatically different results as for the properties of the infinite medium

# Features of the hypernuclear spectroscopy performed through ${}^{A}Z(e, e'K^{+})^{A}_{\Lambda}(Z-1)$ reactions:



Energy calibration Thanks to the availability of Hydrogen targets and hence to the possibility to determine the excitation energy spectrum of the reaction:  ${}^{1}H(e, e'K^{+})\Lambda, \Sigma$  it is possible to obtain very good energy calibration and hence to determine very precisely binding energies



### Energy resolution:

thanks to high resolution spectrometers and the high monochromatic incident electron beam sub-MeV energy resolutions were obtained at Jlab

## Experimental setup and kinematics as originally planned (the experiment was supposed to run in Hall A)



	-			
Beam	$\Delta p/p$	$< 1 \times 10^{-4}~{\rm FWHM}$		
	$E_e$	$4.5~{ m GeV}$		
	D(PCS)	+ QQDQ		
	$\Delta p/p$	$2.6\times10^{-4}~{\rm FWHM}$		
PCS + HRS	$p_{e'}$	$3.0~{\rm GeV}/c\pm4.5\%$		
(e')	$\theta_{ee'}$	$6.5 \pm 1.5 \deg$		
	Solid angle $\Omega_{e'}$	$2.4 \mathrm{msr}$		
	D(PCS)	) + QQD		
	$\Delta p/p$	$4.2  imes 10^{-4}$ FWHM		
	$p_K$	$1.2~{\rm GeV}/c\pm10\%$		
PCS + HKS	$ heta_{eK}$	$12.6\pm4.5~\mathrm{deg}$		
$(K^+)$	Solid angle $\Omega_K$	$7 \mathrm{msr}$		
	Optical length	$12 \mathrm{~m}$		
	$K^+$ survival ratio	26%		





## **Detector setup**



HES

HKS

The detector setup of the experiment E12-20-013 at Jlab is the same as the one employed by the other hypernuclear experiments planned in Hall C (see Toshi's and Sho's talks) but the target used.

### Experiment E12-15-008 at Jlab:

An isospin dependence study of the Lambda-N interaction through the high precision spectroscopy of Lambda hypernuclei with electron beam

Spokesperson:

- F. Garibaldi, T. Gogami, S.N. Nakamura, P. Markowitz, S. Nagao,
- J. Reinhold, L. Tang, G. M. Urciuoli





Ca40 target, 50uA, 100mg/cm2, 230h(9.6days)

## Experiment E12-20-013 at Jlab:

## studying $\Lambda$ interactions in nuclear matter with the $^{208}Pb(e, e'K^+)^{208}_{\Lambda}Tl$ reaction

Spokespersons:

F. Garibaldi, O. Benhar Noccioli, T. Gogami, P. Markowitz, S. Nakamura, J. Reinhold, L. Tang, G. Urciuoli

## The expected spectrum: theoretical calculation



Millener-Motoba calculations performed for the kinematics of the experiment E12-20-013 using the Saclay Lyon elementary amplitudes. The  $\Lambda$  is assumed to be weakly coupled to the proton-hole states of <sup>207</sup>Tl strongly populated in (e,e'p) or (d,3He) reactions on <sup>208</sup>Pb. The  $\Lambda$  single-particle energies were calculated from a Woods-Saxon well fitted to energies derived from the <sup>208</sup>Pb( $\pi^+$ ,  $K^+$ )<sup>208</sup>Pb

## $\Lambda$ interactions in a uniform nuclear medium (1)



The measured charge density distribution of <sup>208</sup>Pb clearly shows that the region of nearly constant density accounts for a very large fraction (~70 %) of the nuclear volume, thus suggesting that its properties largely reflect those of uniform nuclear matter in the neutron star The validity of this conjecture has been long established by a comparison between the results of theoretical calculations and the data extracted from the <sup>208</sup>Pb(e,e'p)<sup>207</sup>Tl cross sections measured at NIKHEF in the 1990s

## $\Lambda$ interactions in a uniform nuclear medium (2)



The energy dependence of the spectroscopic factors extracted from the measured  ${}^{208}Pb(e,e'p){}^{207}Tl$  cross sections compared to the theoretical results shows that short-range correlations appear to be the most important mechanism leading to the observed quenching of the spectroscopic factor, while surface and shell effects only play an important role in the vicinity of the Fermi surface

Deeply bound protons in the <sup>208</sup>Pb ground state are largely unaffected by finite size and shell effects and behave as if they were in nuclear matter

### The expected spectrum: Montecarlo simulations



Expected binding-energy spectrum for the  ${}^{208}Pb(e, e'K^+){}^{208}_{\Lambda}Tl$  reaction with the resolution of 0.6MeV in FWHM. Beam time of 480 hours with the beam intensity of 25 µA is assumed.

Expected binding energy spectrum resolution: 0.6 MeV Expected  $|\Delta B_{\Lambda}| < 100 \text{ keV} (\approx 50 \text{ keV})$ 

## **Experiment Beam Time requests**

### 12-15-008

An isospin dependence study of the Lambda-N interaction through the high precision spectroscopy of Lambda hypernuclei with electron beam

E12-20-013

studying  $\Lambda$  interactions in nuclear matter with the  $^{208}Pb(e, e'K^+)^{208}\Lambda Tl$  reaction

Likely experiment E12-20-013 beam time request has to be increased

Target	Beam	Target	Assumed	Expected	Num. of	Req.	B.G.	S/N	Comments
(Hyper	current	thickness	cross	yield	events	beamtime	rate	-	
Nucleus)	$(\mu A)$	$(mg/cm^2)$	section	(/h)		(hours)	(/MeV/h)		
			(nb/sr)						
$CH_2$	2	500	1000	8.62	1000	120	0.03	290	Calibration
<sup>6</sup> Li	50	150	10	1.51			0.86	1.7	Separate LoI
${}^{9}\mathrm{Be}$	50	150	10	1.01			0.84	1.25	Separate LoI
$^{11}B$	50	150	30	2.47			0.96	2.57	Separate LoI
$^{12}C$	50	150	90	6.79	1100	168	1.20	5.67	Calibration
<sup>27</sup> Al	50	150	60	3.88	330	168	1.06	1.87	Calibration
Subtotal						456			Calibration
${}^{40}Ca~({}^{40}_{\Lambda}K)$	50	150	50	1.13	520	456	2.41	0.47	Physics
${}^{48}Ca~({}^{48}_{\Lambda}K)$	50	150	50	0.94	520	552	1.89	0.50	Physics
Subtotal						1008			Physics
Total						1464			

#### Yields estimation and beam time requirement.

Target and objective	Beam	Target	Assumed	Expected	Num. of	Req.	B.G.	S/N	Comments
hypernucleus	current	thickness	cross	Yield	events	beamtime	Rate	(±4 σ)	
	(µA)	$(mg/cm^2)$	section	(/hour)		(hours)	(/MeV/h)		
			(nb/sr)						
CH <sub>2</sub>	2	500	200	19	1000	54	0.05	252	Calibration
<sup>6,7</sup> Li	50	100	10	5.4	150	28	1.3	4.9	Calibration
°Be	100	100	10	36	300	9	4.7	8.8	Calibration
<sup>10,11</sup> B	25	100	10	16	150	19	0.29	33	Calibration
<sup>12</sup> C	100	100	100	54	2000	37	4.4	17	Calibration
Subtotal for calibration						147			
<sup>208</sup> Pb	25	100	80(g.s.)	0.3	145	480	0.1	21	Production

### LOI12-23-011

High-resolution spectroscopy of light hypernuclei with the decay-pion spectroscopy



- High-resolution & High-precision hypernuclear mass spectroscopy
  - Stopping in a target
  - Two-body decay with π<sup>-</sup> & nucleus
     → hypernuclear ground-state
- > Momentum resolution  $\Delta p \sim 0.1 \text{ MeV/c}$
- > Mass precision  $\Delta M \sim 0.01 \text{ MeV/c}^2$
- Good calibration sources
- Tagging Kaon





#### The expected spectra



- ➤ 300 keV/c (FWHM) peak resolution, <keV precision for <sup>4</sup><sub>Λ</sub>H (stability check is available)
- Yields were estimated by using MAMI results, AMD calc. [Y.Nara's calc. in A.Kawachi Ph.D thesis (U-Tokyo)], and shell-model calc[NPA 489 (1988) 683.].
- Li target: Measurement of s-shell hypernuclei. Benchmark
- Be target: Interesting n-rich hypernuclei
- C target: Additional p-shell hypernuclei

## LOI12-23-013

### Study of charge symmetry breaking in p-shell hypernuclei

determine the energies of ground-state peaks of  ${}_{\Lambda}^{6}$ He,  ${}_{\Lambda}^{9}$ Li, and  ${}_{\Lambda}^{11}$ Be with the accuracy

of  $|\Delta B_{\Lambda}^{\text{total}}| = 70$  keV for the study of CSB.

	Electron-beam experiment at JLab E Emulsion experiment Hadron-beam experiment at J-PARC H $\gamma$ -ray experiment						
H	ypernucleus	T<0	T>0	CSB study			
shell		4H ↓		4He	0	0	0
<u>s</u>					$\square$	$\bigcirc$	
			77.4			~	
		<sub>Λ</sub> He 🕅	<u>∕</u> ,L1* ♥♥♥	<sub>Λ</sub> Be ₩₩	Ο	0	0
ell		<sup>8</sup> Li 😽		<sup>8</sup> <sub>A</sub> Be 🐺	Ο	0	0
hs-q		<sup>9</sup> Li new	βBe ⊮	${}^9_{\Lambda}B$		ightarrow	igodol
		$^{10}_{\Lambda}\mathrm{Be}$		$^{10}_{\Lambda}\mathrm{B}$			igodol
		$^{11}_{\Lambda}\mathrm{Be}_{new}$	$^{11}_{\Lambda}\mathrm{B}$	$^{11}_{\Lambda}C$ VP H		ightarrow	igodol
		$^{12}_{\Lambda}\mathrm{B}$		$^{12}_{\Lambda}\mathrm{C}$			igodol

#### Description of nuclear deformation • Nuclear quadrupole deformation $(\beta, \gamma)$ • β: degree of quadrupole deformation Middle • γ: (tri)axiality ong Short Shor **Triaxial deformation** $\approx 30^{\circ}$ no symmetry axis $\gamma = 60^{\circ}$ Long **Oblate deformation** n ong ß short axis symmetry $= 0^{\circ}$ $\beta = 0$ Short **Spherical Prolate deformation** What is expected in deformed $\Lambda$ hypernuclei

#### Deformation change

 $\bullet\,\Lambda$  particle can change nuclear deformation

#### $\bullet$ Difference of $\mathbf{B}_{\Lambda}$ depending on nuclear deformation

• Energy shifts in excitation spectra

#### $\bullet$ Coupling of $\Lambda$ to deformed nuclei shows unique structure

• For example, rotational band, mixing of configuration, ... etc.

### Description of nuclear deformation

#### •Nuclear quadruple deformation ( $\beta$ , $\gamma$ )



Observing the 3 different *p*-states is strong evidence of triaxial deformation

### LOI Beam Time requests

#### LOI12-23-011

### High-resolution spectroscopy of light hypernuclei with the decay-pion spectroscopy

#### LOI12-23-013

Study of charge symmetry breaking in p-shell hypernuclei

#### LOI12-23-016

Study of a triaxially deformed nucleus using a Lambda particle as a probe

### Parasitically running

Desetion	Assumed cross section	Necessary bear	m time (/days)	Ratio $[t(I_{\rm beam}^{50})/t(I_{\rm beam}^{30})]$		
Reaction	[/(nb/sr)]	$I_{\rm beam}^{30}=30~\mu{\rm A}$	$I_{\rm beam}^{50}=50~\mu{\rm A}$	(×1.7)		
${}^{6}\mathrm{Li}(e,e'K^{+})^{6}_{\Lambda}\mathrm{He}$	10	7	5			
${}^{9}\mathrm{Be}(e,e'K^{+})^{9}_{\Lambda}\mathrm{Li}$	7.6	19.5	14	$\times \frac{1}{1.4} = 0.71$		
${}^{11}\mathrm{B}(e,e'K^+)^{11}_{\Lambda}\mathrm{Be}$	30	3.5	2.5			

Target	Beam	Target	Assumed	Expected	Num. of	Req.	B.G.	S/N	Comments
(Hyper	current	thickness	cross	yield	events	beamtime	rate		
Nucleus)	$(\mu A)$	$(mg/cm^2)$	section	(/h)		(hours)	(/MeV/h)		
			(nb/sr)						
$\operatorname{CH}_2(\Lambda, \Sigma^0)$	2	500	1000	8.62	1000	120	0.03	290	Calibration
${}^{12}C~(^{12}_{\Lambda}B)$	50	150	90	6.79	1100	168	1.20	5.67	Calibration
Subtotal									Combined with
									E12-15-008
$^{27}\text{Al}\left(^{27}_{\Lambda}\text{Mg}\right)$	50	100	$10 \ (p^{\Lambda})$	0.22	150	672	0.78	0.19	Physics
Total						672			

The MOLLER experiment will measure the parity-violating asymmetry  $(A_{PV})$  (that is the fractional difference between rightand left-handed electrons) in the scattering of longitudinally polarized electrons on unpolarized target electrons



Feynman diagrams for Møller experiment at tree level

$$A_{PV} = \frac{\sigma_R - \sigma_L}{\sigma_R + \sigma_L} = mE \frac{G_F}{\sqrt{2}\pi\alpha} \frac{4\sin^2\theta}{(3 + \cos^2\theta)^2} Q_W^e = mE \frac{G_F}{\sqrt{2}\pi\alpha} \frac{2y(1-y)}{1 + y^4 + (1-y)^4} Q_W^e$$

 $\alpha$  fine structure constant; E the incident beam energy; m the electron mass,  $\theta$  the scattering angle in the center of mass frame, y = 1 - E'/E; E' the energy of one of the scattered electrons.  $Q_W^e$  is the weak charge of the electron, proportional to the product of the electron's vector and axial-vector couplings to the Z<sup>0</sup> boson. The electroweak theory prediction at tree level in terms of the weak mixing angle is  $Q_W^e = 1 - 4\sin^2\theta_W$ 



The electroweak theory prediction at tree level for  $Q_W^e$  ( $Q_W^e = 1 - 4\sin^2\theta_W$ ) is modified at the 1-loop level and becomes dependent on the energy scale at which the measurement is carried out ( $\sin^2\theta_W$  runs)".

At MOLLER energy scale,  $Q_W^e \sim 0.0435$ . This is a 42% change of  $Q_W^e$  tree level value of  $\sim 0.075$  (when evaluated at M<sub>z</sub>).

The prediction for the A<sub>PV</sub> measured by MOLLER is 33 parts per billion (ppb). MOLLER goal is to measure this quantity with a precision of 0.7 ppb, achieving a 2.4% precise measurement of  $Q_W^e$ .

The reduction in the numerical value  $Q_W^e$  leads to increased fractional accuracy in the determination of the weak mixing angle of ~ 0.1%

The fact that theoretical uncertainties for the purely leptonic Møller Parity Violation are well under control has to be stressed.

## Search for Physics Beyond the Standard Model (3)





 $\sin^2 \theta_{\rm W}$  vs m<sub>H</sub>. The yellow band is the world average. The black points are the two most precise measurements at  $Q^2 \ll M_Z^2$ . The projected MOLLER error is shown in red.



A variety of BSM dynamics, can have a significant impact on low Q<sup>2</sup> observables while having much reduced impact on corresponding measurements made at colliders. This is because interference effects are highly suppressed on top of the Z<sup>0</sup> resonance. The MOLLER measurement would achieve a sensitivity of  $\delta(\sin^2\theta_W)$ = 0.00028. The projected uncertainty from forward-backward asymmetries after 300 fb<sup>-1</sup> integrated luminosity at the LHC is a systematics limited  $\delta(\sin^2\theta_W)$  = 0.00036, with the dominant error being from parton distribution function (PDF) uncertainties.

Search for Physics Beyond the Standard Model (4) (Complementarity between MOLLER and LHC, first scenario)

- If the LHC observes an anomaly, the MOLLER experiment will provide important constraints to choose among possible BSM scenarios which could to explain the anomaly.
- For example, if the particle predicted by the Minimal Supersymmetric Standard Model were observed, it would be very important to determine if they were generated through radiative loop effects (R-parity conserving) or tree-level interactions (R-parity violating). In fact, RPC would imply that the lightest supersymmetric particle is stable and therefore it would be an obvious candidate for the non-baryonic dark matter which is needed to understand galactic-scale dynamics. On the other hand, RPV would imply that neutrinos are Majorana particles. The RPV and RPC models generate effects of opposite sign in the weak charge, observable by MOLLER.

## Search for Physics Beyond the Standard Model (5)

(Complementarity between MOLLER and LHC, second scenario)

- If the LHC continues to agree with the Standard Model with high luminosity running, the MOLLER experiment, thanks to its measurement at low Q<sup>2</sup>, can still discover signatures of new physics not detectable by LHC but supposed to exist by theoreticians, like:
- Compressed supersymmetry
- Lepton number violating amplitudes mediated by doubly charged scalars → extended Higgs sector models containing complex triplet representations of SU(2).
- And ....

## Search for Physics Beyond the Standard Model (6) (Complementarity between MOLLER and LHC, second scenario, continued)



Dark photon search.

Experiments measuring Parity Violating Electron Scattering (PVES) improved constantly their techniques. This allows us, today, to measure extremely small Parity violating Asymmetry (APV)

Pioneering (1978) early SM tests SLAC E122 PVDIS – Prescott *et al.* A = -152 ppm Bates <sup>12</sup>C, Mainz Be

Strange Form Factors (1998 –2009) SAMPLE, HAPPEX, G0, A4  $A \sim 1 - 50 \text{ ppm}$ 

#### Standard Model Tests (2003 - 2018)

SLAC E158 Moller: A = - 131 ppb (13% precision on electron's weak charge) JLAB Qweak: A = -230 ppb

#### Neutron radii: (2012-2022)

JLab: PREX-I, PREX-II, CREX

Future: Standard Model, hadron structure studies: MOLLER: A = - 35 ppb Goal: 2.5% precision on electron's weak charge



#### JLab E12-09-005: The Measurement of a Lepton-Lepton **Electroweak Reaction (MOLLER)** e-

Parameter	Value			
E [GeV]	$\approx 11.0$			
E' [GeV]	2.0 - 9.0			
$\theta_{cm}$	50°-130°			
$ heta_{lab}$	0.26°-1.2°			
$< Q^2 > [ m GeV^2]$	0.0058			
Maximum Current [ $\mu$ A]	70			
Target Length (cm)	125			
$\rho_{tgt} \text{ [g/cm^3]} (\text{T}= 20\text{K}, \text{P} = 35 \text{ psia})$	0.0715			
Max. Luminosity $[cm^{-2} sec^{-1}]$	$2.4 \cdot 10^{39}$			
$\sigma$ [ $\mu$ Bam]	$\approx 60$			
Møller Rate @ 65 µA [GHz]	$\approx 134$			
Statistical Width(1.92 kHz flip) [ppm/pair]	$\approx 91$			
Target Raster Size [mm]	5 x 5			
Production running time	344 PAC-days = 8256 hours			
$\Delta A_{raw}$ [ppb]	$\approx 0.54$			
Background Fraction	$\approx 0.10$			
$P_{beam}$	$\approx 90\%$			
$< A_{pv} > [ppb]$	$\approx 32$			
$\Delta A_{stat} / < A_{expt} >$	2.1%			
$\delta(\sin^2 \theta_W)_{stat}$	0.00023			





- Integrate counts over each helicity state at 1.92 kHz
- Yields normalized to the Beam Charge Monitors
- counting statistical width = 91 ppm , to keep the beam ٠ intensity measurement an insignificant source of noise: BCM resolution ~ 10 ppm

MOLLER CDR

### Experimental apparatus





