

## Oxygen fugacity vs. mineralogical control on transition metal (Fe, Cr, V) stable isotope compositions of Mariana forearc peridotites

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Controversy surrounds the oxygen fugacity ( $fO_2$ ) of subduction zones. Emerging redox-sensitive stable isotope systems may provide an independent assessment if their isotopic fractionation can be linked to  $fO_2$ . However, other factors such as mineral coordination number can also influence stable isotope fractionation. Here we present the first investigation of combined Fe, Cr and V stable isotope compositions. We investigate peridotites from two forearc seamounts drilled on ODP Leg 125. The studied peridotites have been characterised for major and trace elements, modal mineralogy and  $fO_2$  (ranging from FMQ -0.7 to +1.8) [1]. We find no correlation between Fe, Cr, or V isotopes and  $fO_2$ . Iron isotope compositions are generally heavier than the terrestrial mantle [e.g., 2], but show significant scatter. Chromium and V isotopes are positively correlated with V isotopes displaced to heavier values than comparably depleted peridotites [3]. Chromium isotope compositions are within the range published for mantle xenoliths [4]. However, Cr isotopes correlate with modal clinopyroxene in forearc harzburgites. These results suggest that Cr and V isotopes may be more robust to secondary processes than Fe isotopes and that mineralogy may have a greater influence on Cr and V stable isotope fractionation than oxygen fugacity.

[1] Parkinson and Pearce, 1998. *JPet*, **39**, 1577-1618. [2] Craddock *et al* 2013. *EPSL*, **365**, 63-76. [3] Prytulak *et al* 2013. *EPSL*, **365**, 177-189. [4] Schoenberg *et al* 2008. *Chem. Geol.*, **249**, 294-306.

## Radiogenic heat potential of the Sardinian Variscan crust

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The <sup>40</sup>K, <sup>238,235</sup>U and <sup>232</sup>Th composition of the Variscan crust is derived from *in-situ* radiological characterization using a portable gamma-ray spectrometer [1]. Details about the analytical procedure and statistical processing of spectrometric data is given in [2]. A total of 400 measurements were performed in Sardinia, because of excellent exposure and detailed geologic information on the architecture of the Corsica-Sardinia Batholith and its country rocks [3].

The results of gamma-ray spectrometry measurements indicate that most granitoids have potential heat production rate between 0.49 and 6.92  $\mu\text{W m}^{-3}$ . Both migmatites and low-grade metasediments are characterized by lower values in the range 1.3 – 3.1  $\mu\text{W m}^{-3}$ . The U-Th-K abundances in the Sardinian Variscan crust are  $1283^{+340}_{-463}$  Bq kg<sup>-1</sup>,  $47^{+18}_{-29}$  Bq kg<sup>-1</sup> and  $67^{+21}_{-31}$  Bq kg<sup>-1</sup> respectively for <sup>40</sup>K, <sup>238</sup>U and <sup>232</sup>Th. These values are slightly higher than those typical for the upper continental crust [4]. However, the average heat production rate of the Variscan crust of northern Sardinia is about half of that inferred in the Bohemian Massif [5]. Therefore, we argue that selective enrichment of heat-producing elements in the crust cannot account for early Permian HT metamorphism in this part of the Variscan chain.

[1] Caciolli *et al* (2012) *Sci. Total Environ* **414**, 639–645. [2] Puccini *et al* (2013) *Environ. Earth Sci* in press, doi: **10.1007/s12665-013-2442-8**. [3] Casini *et al* (2012) *Tectonophysics* **544**, 31–49. [4] Rudnick & Gao (2003) *Treatise on geochemistry: meteorites, comets and planets*, vol 1, Elsevier Ltd., Oxford. [5] Lexa *et al* (2011) *J. Metamorphic Geol* **29**, 79–102